

Darmstadt 2008

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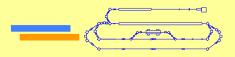
α-Cluster States in Electron Scattering *

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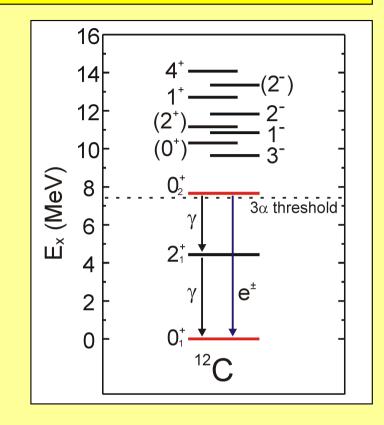
* Supported by DFG under contract SFB 634





Motivation: structure of the Hoyle state

- Hoyle state is a prototype of α-cluster states in light nuclei
- Cannot be described by shell-model approaches
- α-cluster models predict Hoyle state as a dilute gas of weakly interacting α particles resembling the properties of a Bose-Einstein Condensate (BEC)

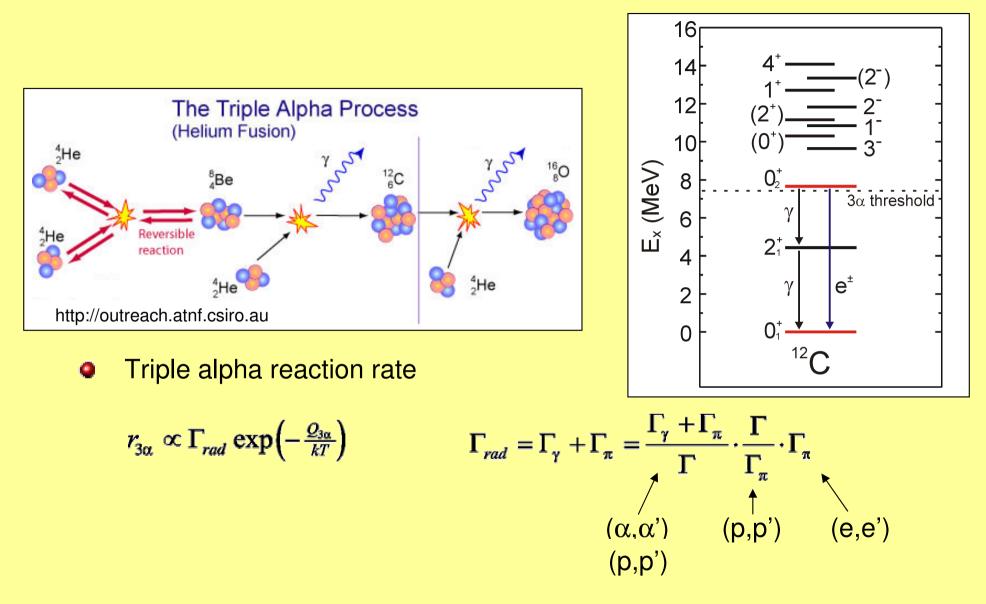


 Comparison of high-precision electron scattering data with predictions of FMD and α-cluster models

Hoyle state cannot be understood as a true Bose-Einstein Condensate !

 M. Chernykh, H. Feldmeier, T. Neff, P. von Neumann-Cosel, and A. Richter, Phys. Rev. Lett. 98 (2007) 032501

Motivation: astrophysical importance



Reaction rate with accuracy ~ 6% needed

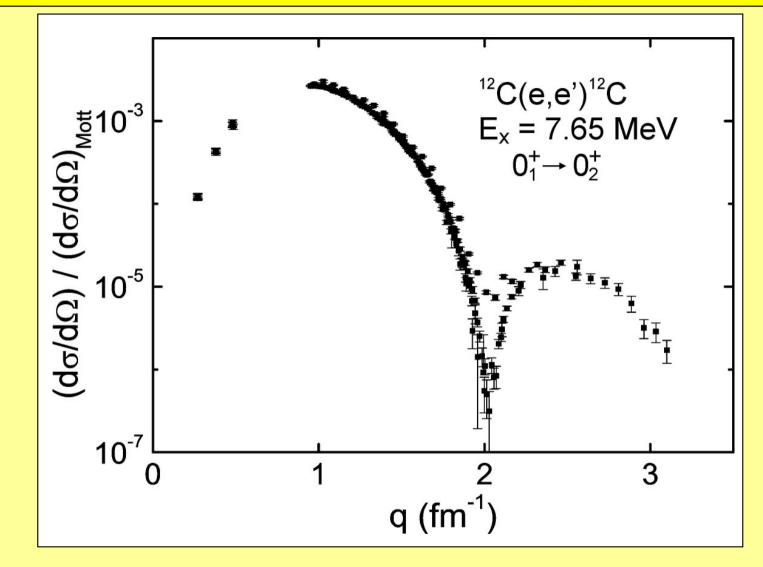
S.M. Austin, Nucl. Phys A758 (2005) 375c

Motivation: astrophysical importance

$r_{3\alpha} \propto \Gamma_{rad} \exp$	$\Gamma_{rad} = \Gamma_{\gamma} + 1$	$\Gamma_{\pi} = \frac{\Gamma_{\gamma} + \Gamma_{\pi}}{\Gamma} \cdot \frac{\Gamma}{\Gamma_{\pi}} \cdot \Gamma_{\pi}$
Quantity	Value	Error (%)
Q_{3lpha}	$379.38\pm0.20~{\rm keV}$	$1.2 \ (T_9 = 0.2)$
Γ_{rad}/Γ	$(4.12 \pm 0.11) \times 10^{-4}$	2.7
Γ_π/Γ	$(6.74 \pm 0.62) \times 10^{-6}$	9.2
Γ_{π}	$(62.0 \pm 6.0) \times 10^{-6} \text{ eV}$	9.7 Crannell <i>et al.</i> (1967)
Γ_π	$(59.4 \pm 5.1) \times 10^{-6} \text{ eV}$	8.6 Strehl (1970)
Γ_{π}	$(52.0 \pm 1.4) \times 10^{-6} \text{ eV}$	2.7 Crannell <i>et al.</i> (2005)

• Total uncertainty $\Delta r_{3\alpha}/r_{3\alpha} = 11.6\%$ only

Transition form factor to the Hoyle state



• Fourier-Bessel analysis: Crannell (2005)

Extrapolation to zero momentum transfer: Crannell (1967), Strehl (1970)

H. Crannell, data compilation

Model-independent PWBA analysis

$$\left(\frac{d\sigma}{d\Omega}\right)_{PWBA} = 4\pi \left(\frac{e^2}{E_0}\right)^2 f_{rec} \ V_L(\theta) \ B(C0,q)$$

$$4\pi B(C0,q) = \left[\langle 0_2^+ | \int \hat{\rho}_N j_0(qr) d^3r | 0_1^+ \rangle\right]^2$$

$$\langle r^\lambda \rangle_{tr} = \langle 0_2^+ | \int \hat{\rho}_N r^\lambda d^3r | 0_1^+ \rangle$$

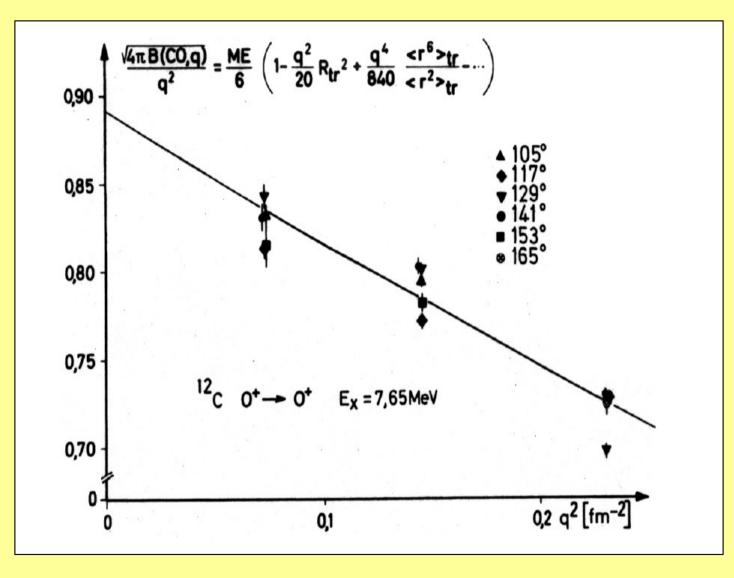
$$ME = \langle r^2 \rangle_{tr}, \qquad R_{tr}^2 = \frac{\langle r^4 \rangle_{tr}}{\langle r^2 \rangle_{tr}}$$

$$\sqrt{4\pi B(C0,q)} = \frac{q^2}{6} (ME) \left[1 - \frac{q^2}{20} R_{tr}^2 + \cdots\right]$$

$$\Gamma_\pi \propto ME^2$$

• Model-independent extraction of the partial pair width Γ_{π}

Monopole matrix element



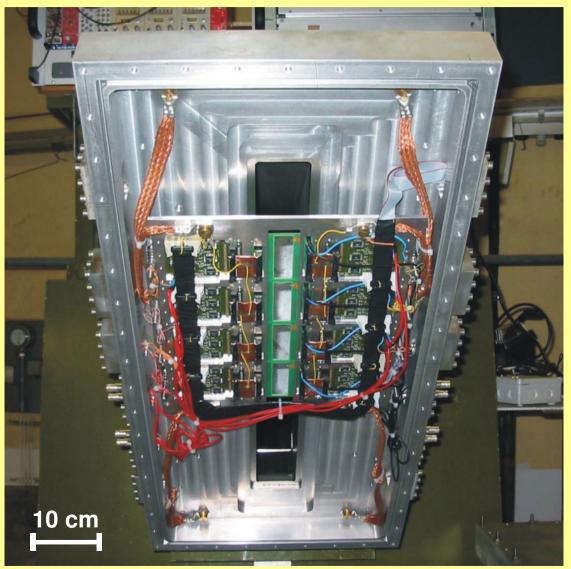
• $ME = 5.37(22) \text{ fm}^2$, $R_{tr} = 4.24(30) \text{ fm}$

Large uncertainty because of narrow momentum transfer region
 P. Strehl, Z. Phys. 234 (1970) 416

Lintott spectrometer

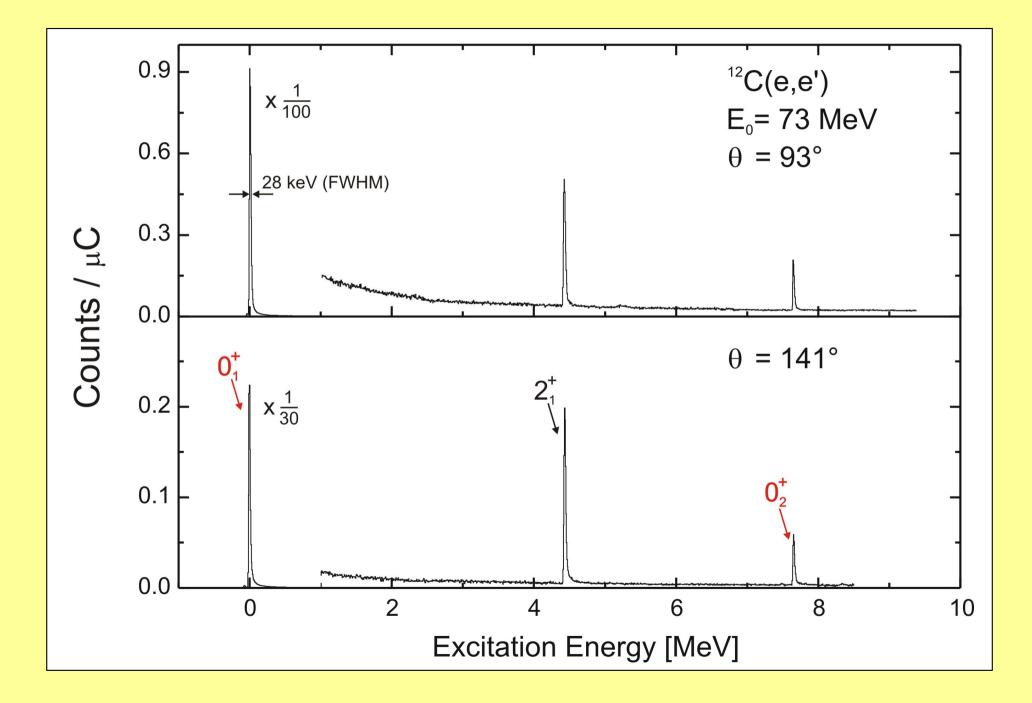


Detector system

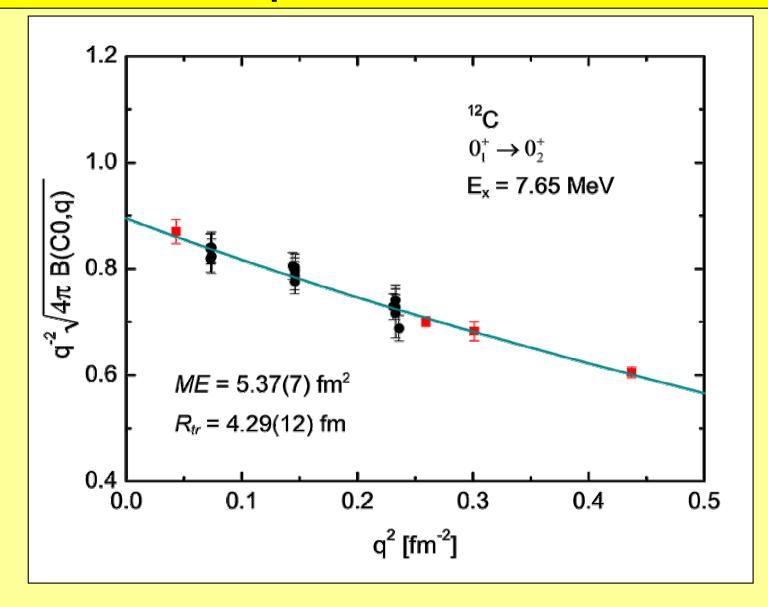


- Si microstrip detector system: 4 modules, each 96 strips with pitch of 650 μm
- Count rate up to 100 kHz
- Energy resolution 1.5x10⁻⁴

Measured spectra



Monopole matrix element



$$\sqrt{4\pi B(C0,q)} = \frac{q^2}{6} (ME) \left[1 - \frac{q^2}{20} R_{tr}^2 + \cdots \right]$$

Triple alpha reaction rate

Quantity	Value	Error (%)	
Q_{3lpha}	$379.38\pm0.20~{\rm keV}$	$1.2 (T_9 = 0.2)$	
Γ_{rad}/Γ	$(4.12 \pm 0.11) imes 10^{-4}$	2.7	
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Γ_{π}	$(52.0 \pm 1.4) \times 10^{-6} \text{ eV}$	2.7 Crannell <i>et al.</i> (2008	5)
Γ_{π}	$(59.6 \pm 1.5) \times 10^{-6} \text{ eV}$	2.5 Present work	

- Total uncertainty $\Delta r_{3\alpha}/r_{3\alpha} = 10\%$
- Only Γ_{π}/Γ need to be improved

Summary and outlook

- Hoyle state is important for stellar nucleosynthesis
- Monopole matrix element can be extracted by extrapolation of cross section to zero momentum transfer
- Γ_{π} for decay of the Hoyle state with uncertainty 2.5% extracted
- Outlook
 - Hoyle state: independent Fourier-Bessel analysis
 - ¹⁶O: broad 0+ state at 15 MeV

Thank you!

Outline

Motivation:

- Astrophysical importance
- Model-independent PWBA analysis
- High-resolution electron scattering measurements
- Results
 - Extraction of monopole matrix element ME
 - Comparison with FMD and α -cluster model predictions
- Summary and outlook

Model-independent PWBA analysis

$$\left(\frac{d\sigma}{d\Omega}\right)_{\rho_{\text{WBA}}} = 4\pi \left(\frac{e^2}{E_0}\right)^2 f_{rec} \quad V_L(\theta) \quad B(C0,q)$$

$$4\pi B(C0,q) = \left[\langle 0_2^+ | \int \hat{\rho}_N j_0(qr) \, d^3r | 0_1^+ \rangle\right]^2$$

$$\langle r^\lambda \rangle_{tr} = \langle 0_2^+ | \int \hat{\rho}_N \, r^\lambda \, d^3r | 0_1^+ \rangle$$

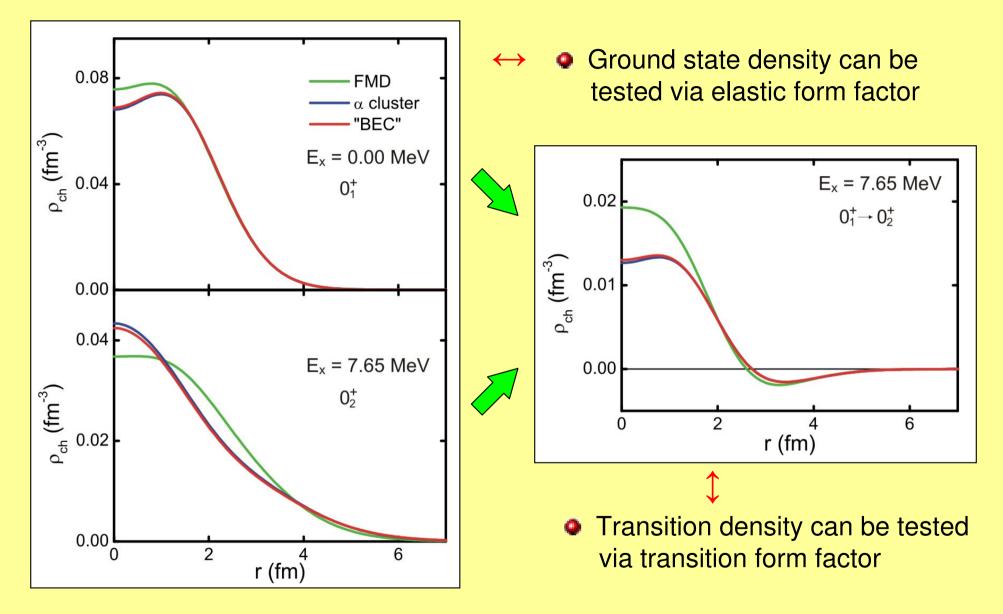
$$ME = \langle r^2 \rangle_{tr}, \qquad R_{tr}^2 = \frac{\langle r^4 \rangle_{tr}}{\langle r^2 \rangle_{tr}}$$

$$\sqrt{4\pi B(C0,q)} = \frac{q^2}{6} (ME) \left[1 - \frac{q^2}{20} R_{tr}^2 + \frac{q^4}{840} \frac{\langle r^6 \rangle_{tr}}{\langle r^2 \rangle_{tr}} - \frac{q^6}{60480} \frac{\langle r^8 \rangle_{tr}}{\langle r^2 \rangle_{tr}} + \cdots \right]$$

$$\frac{\langle r^6 \rangle_{tr}}{\langle r^2 \rangle_{tr}} = x_1 (R_{tr}^2)^2, \quad \frac{\langle r^8 \rangle_{tr}}{\langle r^2 \rangle_{tr}} = x_2 (R_{tr}^2)^3, \quad \cdots$$

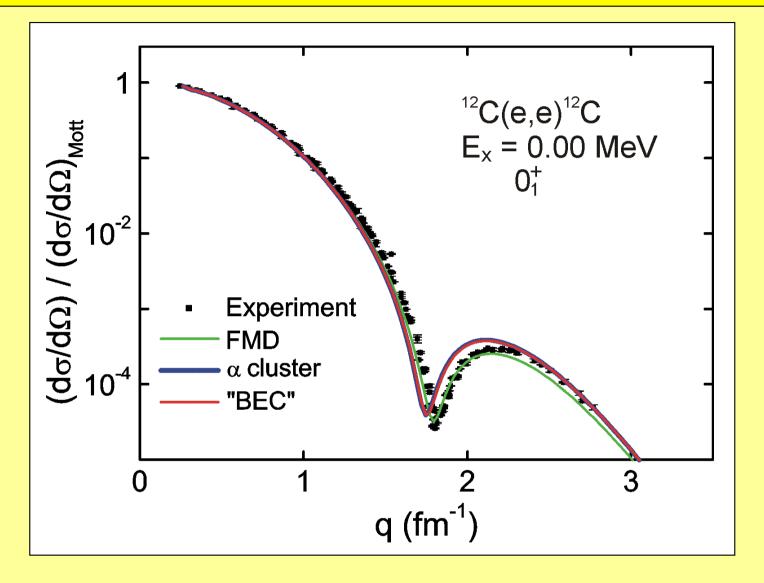
Model-independent extraction of monopole matrix element ME

¹²C densities



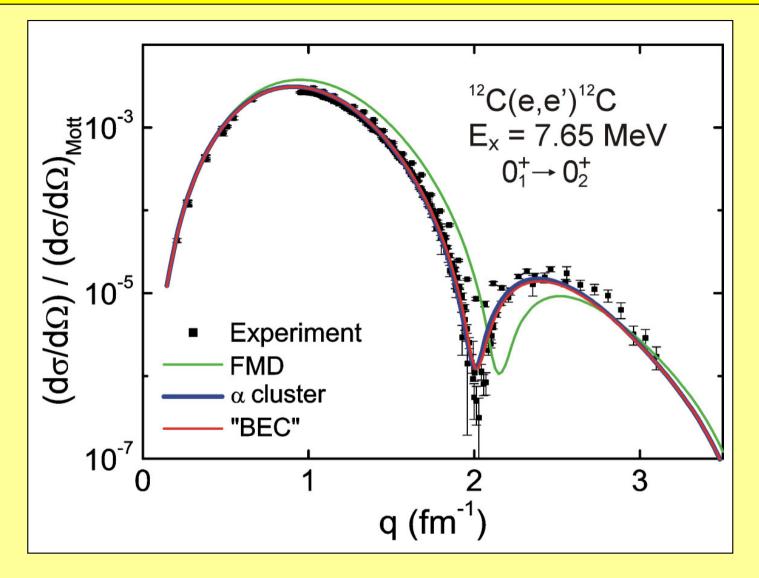
FMD : R. Roth, T. Neff, H. Hergert, and H. Feldmeier, Nucl. Phys. **A745** (2004) 3 "BEC": Y. Funaki *et al.*, Phys. Rev. C **67** (2003) 051306(R)

Elastic form factor



Described well by FMD

Transition form factor to the Hoyle state



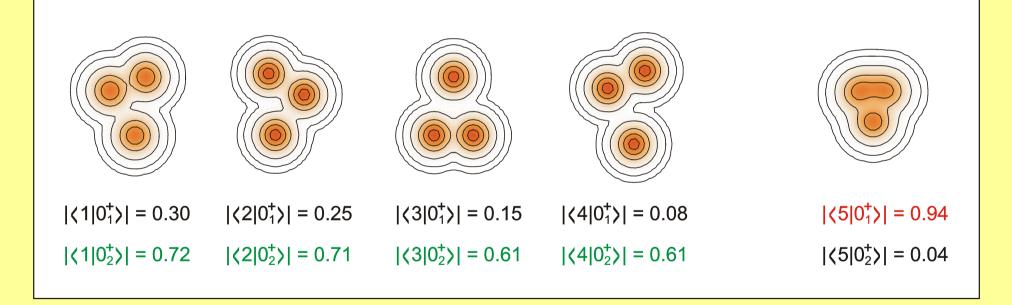
• H. Crannell, data compilation

Described better by α-cluster models

What is actual structure of the Hoyle state ?

 In the "BEC" model the relative positions of α clusters should be uncorrelated

Overlap with FMD basis states



 But in the FMD and α-cluster model the leading components of the Hoyle state are cluster-like and resemble ⁸Be + ⁴He configurations

Summary and outlook

Summary

- Γ_{π} for decay of the Hoyle state with uncertainty 2.5% extracted
- Hoyle state is not a true Bose-Einstein condensate
- ⁸Be + α structure
- Outlook
 - Hoyle state: Fourier-Bessel analysis of all available data
 - ¹²C: 0⁺₃ and 2⁺₂ states
 - ¹⁶O: broad 0⁺ state at 15 MeV