



# LECTURE 3: THE NUCLEAR SHELL STRUCTURE

### NUCLEAR STRUCTURE STUDIED WITH SPECTROSCOPY AND REACTIONS

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TU Darmstadt, IKP, February 2017



- The shell model and shell closures
- Origins of shell evolution
  - diffusiveness of the nuclear surface
  - spin-isospin and tensor terms of the NN interaction
  - reduction of spin-orbit
  - 3N forces
- RI beam production and future facilities
- The N=16 "new magic number", collapse of N=20, 28 shell closures
- Mirror region: the N=40 island of inversion and the <sup>78</sup>Ni region
- Are N=32 and 34 "new" magic numbers?
- Heavier doubly-magic nuclei: <sup>100</sup>Sn and <sup>132</sup>Sn

## 22 The shell model: history

- **1932**: Chadwick discovers the neutron (Nobel 35')
- W.M. Elsasser, J. Phys. Radium 4, 549 (1933)
  Magic numbers from binding energies, abundancies
- M. Goeppert-Mayer, Phys. Rev. 74, 235 (1948)
  « On closed shells in nuclei »
- □ H.E. Suess, J.H.D Jensen, Arkiv. Fys. 3, 577 (1951)
- 1963: Nobel prize in Physics to Mayer and Jensen for « their discoveries concerning nuclear shell structure »



M. Goeppert-Mayer



er J.H.D. Jensen



E. Wigner\*

 E. Wigner won ½ of the 1963 Nobel Prize for his contribution to the theory of the atomic nucleus and fundamental symetries



### Cea Evidence for shells

- □ Several observables show a pattern consistent with the simple shell model picture
- □ Abundances, separation energy, first excitation energy, moments, spins,...
- ❑ Huge efforts over the past decades in experimental nuclear physics

#### W.M. ELSASSER

#### La structure des noyaux atomiques complexes

Annales de l'I. H. P., tome 5, nº 3 (1935), p. 223-262.



First 2<sup>+</sup> excitation energies (even-even nuclei in 2016)



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$$\langle j^{2}(J=0)j' | V_{1n} + V_{2n} | j^{2}(J=0)j' \rangle = 2 \sum_{\substack{J=|j-j'|\\ J=|j-j'|}} (2J+1) \langle jj'J | V | jj'J \rangle / \sum_{\substack{J=|j-j'|\\ J=|j-j'|}} (2J+1).$$
(1)

### Cell First evidence for non universality of magic numbers

□ Mass measurement of <sup>26-32</sup>Na leads to an unexpected region of deformation at N=20

C. Thibault et al., Phys. Rev. C 12 (1975)

Then the behavior of the experimental data for the sodium isotopes at N = 20 is inconsistent with the classic shell closure effect, and more reminiscent of the behavior one observes when entering a region of sudden deformation.

□ Low 2<sup>+</sup> excitation energy (885 keV) of <sup>32</sup>Mg

C. Détraz et al., Phys. Rev. C 19, 164 (1979)

(ii) The confirmation of the onset of the new region of deformation at  $Z \simeq 11$ and N = 20 definitively observed with spectroscopic tools. The location of the first 2<sup>+</sup> state of <sup>32</sup>Mg at an energy of only 885 keV (fig. 24) is associated with a large deformation at N = 20.

 $\Box \quad \text{High BE2 value for 0+->2+ in } ^{32}\text{Mg}$ 

T. Motobayashi et al., PLB 346, 9 (1995)





# Cea Shell evolution with isospin



shell structure is not yet known experimentally nor fully understood theoretically
 main reason to investigate radioactive nuclei

stable=specific r

radioactive=general

experimental difficulty : low production rates for most exotic nuclei
 new methods and priviledged observables suited for low rates

### Shell evolution at the drip line

VOLUME 72, NUMBER 7

#### PHYSICAL REVIEW LETTERS

**14 FEBRUARY 1994** 

Nuclear Shell Structure at Particle Drip Lines

J. Dobaczewski,\* I. Hamamoto,<sup>†</sup> W. Nazarewicz,\* and J. A. Sheikh<sup>‡</sup>

Shell evolution investigated at the HF level

□ shell structure of neutron dripline nuclei strongly affected by interaction with continuum

Significant difference between HF and RMF predictions for isospin dependence for the neutron spin-orbit splitting

□ Large diffusiveness of neutron density and central potential lead to a single-particle spectrum close to that of the **harmonic oscillator with spin-orbit term** 

conclusion do not apply to most studied medium-mass nuclei still quite far from the neutron dripline



# Cea Spin-orbit one-body potential

 One-body spin-orbit interaction depends on the derivative of the nuclear density

$$V_{\ell s}^{n}(r) \propto \frac{1}{r} \frac{\partial}{\partial r} \left[ 2\rho_{n}(r) + \rho_{p}(r) \right] \vec{\ell}.\vec{s}$$
$$V_{\ell s}^{p}(r) \propto \frac{1}{r} \frac{\partial}{\partial r} \left[ \rho_{n}(r) + 2\rho_{p}(r) \right] \vec{\ell}.\vec{s}$$

D. Vautherin, D.M. Brink, Phys. Rev. C5 (1972) 626.

- weighting factors in front of p,n densities depend on the formalism
- Diffuse-surface nuclei, for example at the neutron-dripline, may experience a smaller spin-orbit splitting



# Origin of shell evolution: monopole drift

□ Hamiltonian can be decomposed in a monopole and a multipole part

 $\mathcal{H} = \mathcal{H}_m$ (monopole) +  $\mathcal{H}_M$ (multipole).

**monopole**: effective single-particle energies  $\langle SC \pm 1 | H | SC \pm 1 \rangle = \langle SC \pm 1 | H_m | SC \pm 1 \rangle$ SC: Shell Closure (core or filled valence space)

**multipole**: correlations (Qpole, pairing,...)

Gaps evolve linearly with the number of nucleons when an orbital is filled

□ Changes depend on the 2B matrix elements

*Example:* neutron gap between orbitals  $n_1, n_2$  with proton number varying from 0 to  $(2j_{p1}+1)$ 

 $\Delta \varepsilon_{n1} = x V_{j_{p1}j_{n1}}^{pn}$  $\Delta (gap) = (2j_{p1} + 1)(V_{j_{p1}j_{n1}}^{pn} - V_{j_{p1}j_{n2}}^{pn})$ 



Figure from

O. Sorlin, M.-G. Porquet, Prog. In Part. Nucl. Phys. 61, 602 (2008)

## Cea The example of proton "ESPE" in Sb isotopes



J.P. Schiffer et al., PRL 92 (2004)



Central part of NN force depends on the relative distance with spin and isospin dependences The **spin-isospin term** belongs to the central part of the effective interaction: 

$$V_{\tau\sigma} = \tau . \tau \sigma . \sigma f_{\sigma\tau}(r)$$









T. Otsuka et al., PRL 87, 082502 (2001)





The tensor force contains an angular momentum, spin and isospin component:

$$V = (\overrightarrow{\tau_1}.\overrightarrow{\tau_2})(\left[\overrightarrow{\sigma_1}.\overrightarrow{\sigma_2}\right]^{(2)}.Y^{(2)})f(r)$$

proton-neutron interaction with opposite spin is attractive and claimed to play a major role in shell evolution
 T. Otsuka *et al.*, PRL **95**, 232502 (2005)



Extension of the spin-isospin mechanism for proton and neutron in different shells

EDF approaches with Skyrme or Gogny forcesdo not include explicite tensor term

Explicit introduction of tensor term in effective
 Skyrme functionals does not give obvious
 improvement of spectroscopy
 T. Lesinski *et al.*, PRC **76**, 014312 (2007)



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### Cea Three-body forces

 ❑ Nucleons are not point like particles
 ❑ Some internal degrees of freedom are neglected Example: Nucleons may excite into a ∆ resonnance

 $\begin{bmatrix} 2N & 3N & 4N \\ X & H & - & - \\ M & X & H & - & - \\ NLO & X & H & - & - \\ M & - & - & - \\ M$ 

#### pol(p)-d elastic scattering analysing power



K. Sekiguchi et al., PRC 83, 061001 (2011)

#### **Chiral Effective Field Theory** (EFT):

- pion is the only explicit degree of freedom
- expansion in powers of momenta
- many-body forces emerge naturally

### Cal Shell evolution: 3N forces

PRL 105, 032501 (2010)

PHYSICAL REVIEW LETTERS

week ending 16 JULY 2010

#### **Three-Body Forces and the Limit of Oxygen Isotopes**

Takaharu Otsuka,<sup>1,2,3</sup> Toshio Suzuki,<sup>4</sup> Jason D. Holt,<sup>5</sup> Achim Schwenk,<sup>5</sup> and Yoshinori Akaishi<sup>6</sup>

- □ 3N forces explain the dripline location for oxygens
- 3N forces between 2 valence neutrons and one core neutron induce repulsion among excess neutrons (linked to Pauli exclusion principle)



- □ Modify the ESPE (monopole) of valence orbitals
- □ Key to explain the magic character of <sup>48</sup>Ca J. Holt et al., J. Phys. G 39, 085111 (2012)



### **Spectroscopic observables and shell closures**

- Masses
- □ beta-decay lifetime
- Excitation energy of the first excited state (most of the time 2<sup>+</sup> in even-even nuclei)
- **Transition probabilities**  $B(E2;0^+->2^+)$

#### Spectroscopy:

- □ spectroscopy is often the first information
- $\begin{tabular}{ll} $\square$ in-beam $\gamma$ and spectroscopy following $\beta$ decay are both competitive $$ $\end{tabular}$



- Spectroscopic factors extracted from direct reaction cross sections
- □ Still uncomplete/inconsistent theory for quantitative interpretation (see Lecture 4, tomorrow)



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### Certain Network Contractive Ion beam production

#### In-Flight Thin target



#### **Chacteristics:**

- □ > 50 MeV/nucleon
- no chemical selection
- □ fast production (short lives < ms)
- limited beam « quality »



#### **Characteristics:**

- □ Low energy (from kV)
- Need for re-acceleration for reactions
- □ high-quality emittance
- □ Chemical selectivity
- □ diffusion / charge breeding time (> 1ms)

## Cer Overview of Radioactive Ion beam production



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### Ceal New ISOL / re-accelerated RI beam facilities worldwide





Proton beam (1.4 GeV, ~2 mA)

Target UC<sub>x</sub> (~ 50 g/cm<sup>2</sup>)

Ionization Laser Selectivity in Z

Magnetic spectrometer Selectivity in A/Q

Post-acceleration (REX & HIE-ISOLDE)









#### □ LINAC cavities: 6MV/m

- September 10<sup>th</sup>, 2016: Miniball started taking data with <sup>110</sup>Sn<sup>26+</sup> at 4.5 MeV/u with 2 CMs for the first time. (stage 1) Up to 5.5 MeV/u for A/Q=4.5 in Spring 2017.
- □ 2017: extension to 10 MeV/nucleon (stage 2)



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- □ RIB production via p-induced fission on UCx target, 10<sup>13</sup> fissions/s
- **□** Proton driver: 35-70 MeV cyclotron, 750 µA
- □ RIB delivered to the existing experimental areas
- □ Status: cyclotron installed, ISOL facility under test

#### **EXPECTED RIB PRODUCTION RATE**





- Next generation European ISOL facility (at the concept stage)
- Superconducting linear accelerator, protons at 1 GeV and 5 MW (and ions up to A=40)
- RIB production mechanism: proton induced AND conversion+neutron induced fission
- Low energy lines AND **reacceleration up to 150 MeV/u**, fragmentation of ISOL beams
- EURISOL DF charged of preparing physics cases and coordinate existing European ISOL facilities



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# Cea Leading fragmentation RIB facilities in the world



### Radioactive Isotope Beam Factory (RIBF)



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### Radioactive Isotope Beam Factory (RIBF)





- Super- conducting double-synchrotron SIS100/300 with magnetic rigidities of 100 and 300 Tm
- Unprecedented variety of accelerated particles from antiprotons to Uranium
- □ Injection from UNILAC and SIS18
- □ Operation expected from >2023
- Different physics communities:

APPA: Atomic Physics, Plasma Physics, Appl.
 CBM: Compressed Baryonic Matter
 NuSTAR: Nuclear Structure & Astrophysics
 PANDA: Hadron Structure & Dynamics





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- next generation RIB facility in the US, first beams expected in 2021
- □ 200 MeV/u linac primary beams (future extension to 400 MeV/u possible)
- Primary beam intensities of 400 kW (<sup>238</sup>U@200 MeV/u = 8 pmA)
- □ Intermediate and low energy:
  - in-flight fission and fragmentation: secondary beams at about 150 MeV/u
  - stopped and reaccelerated beam (up to 12 MeV/u)







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# Cea The N=20 island of inversion (IoI)

Island of Inversion: Due to quadrupole deformation, the « intruder » configuration (np-nh) is energetically favored relative to the « normal » configuration.

#### Mechanism proposed by

- X. Campi et al., Nucl. Phys. A 251, 193 (1975)
- A. Poves, J. Retamosa, Phys. Lett. B 184, 311 (1987)
- W. Warburton, J.A. Becker, B.A. Brown, Phys. Rev. C 41, 1147 (1990).



F. Nowacki, A. Poves, Phys. Rev. C **90**, 014302 (2014)

## Cea Search for 0p-0h 0<sup>+</sup> states at N=20

- □ Competition of « normal » and « intruder » configurations at low excitation energy
- □ How does the inversion occurs: sharp or progressive?
- □ Need to know the relative energy of the 0p-0h and 2p-2h configurations
- **D** Before 2010: the second  $0^+$  state in <sup>34</sup>Si and <sup>32</sup>Mg were not know.


# **Discovery of the 0**<sup>+</sup><sub>2</sub> state in <sup>34</sup>Si at GANIL

All previous experiments (inelastic scattering, beta decay) failed in finding this 0<sup>+</sup><sub>2</sub> state
 0<sup>+</sup><sub>2</sub> expected to be located below the 2<sup>+</sup><sub>1</sub> state and therefore forbidden to decay via gamma
 Should then decay via electron conversion (E<1022 keV) or e<sup>+</sup>e<sup>-</sup> pair creation



# Discovery of the 0<sup>+</sup><sub>2</sub> state in <sup>34</sup>Si at GANIL



# **Discovery of the 0**<sup>+</sup><sub>2</sub> state in <sup>34</sup>Si at GANIL

- Shell model (SDP-U-mix interaction) and experiment in very good agreement
- □ Second 0<sup>+</sup> state in <sup>34</sup>Si is calculated to be 86% 2p-2h
- □ 2713 keV is still a rather high value of excitation energy (sign for a sharp transition to IoI)



### **Second 0<sup>+</sup> of <sup>32</sup>Mg from two-neutron transfer**

### **T-REX+MINIBALL** setup at **ISOLDE, CERN**: particle-γ detection for direct reactions



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... other ongoing developments (TRACE - GASPARD for SPES/GANIL in Europe)

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### **Second 0<sup>+</sup> of <sup>32</sup>Mg from two-neutron transfer**



- □ Ti foil loaded with tritium <sup>3</sup>H/Ti ratio of 1.5
- **□** Equivalent to 40 µg.cm-2 of <sup>3</sup>H
- □ Activity of the target is 10 GBq
- missing mass (t,p): absolute excitation energy
- modest energy resolution (about 800 keV FWHM)
- □  $0_{2}^{+}$  state assigned from angular distribution □  $E_{x}$ =1083(33) keV

K. Wimmer et al., Phys. Rev. Lett. 105, 252501 (2010)



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### **Second 0<sup>+</sup> of <sup>32</sup>Mg from two-neutron transfer**



K. Wimmer et al., Phys. Rev. Lett. 105, 252501 (2010)



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# Cea N=16, a new magic number

- □ Disapearance of the N=20 shell closure and emergence of N=16 connected
- $\Box$  Spin-isospin mechanism between the  $\pi d_{5/2}$  and  $\nu d_{3/2}$
- □ The dripline nucleus <sup>24</sup>O would then be doubly magic
- □ No bound excited state seen in <sup>24</sup>O in in-beam  $\gamma$  spectroscopy: unbound 2<sup>+</sup> (E>S<sub>n</sub>=4.09 MeV)

M. Stanoiu et al., Phys. Rev. C 69, 0234312 (2004).



# Cea N=16, a new magic number



⇒ N=16 ? Ozawa *et al.,* Phys. Rev. Lett. **84**, 5493 (2000)

# Cea N=16, a new magic number



# N=16, N=20: joint spin-isospin mechanism

- □ Intrusion of negative-parity states in neutron-rich N=17 isotones
- □ « measure » of the lowering of the *sd-fp* shell gap



J.R. Terry *et al.*, Phys. Lett. B **640**, 86 (2006) S.M. Brown *et al.*, Phys. Rev. C **85**, 011302(R) (2012)

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### Evolution of the N=28 spin-orbit shell closure

Shell model, 1997

Shell gap at N=28 smaller in Si than in Ca

- <sup>42</sup>Si still predicted to be magic
- Shell model, 2004

E. Caurier, F. Nowacki and A. Poves, NPA742(2004) Adjustment of interaction on <sup>35</sup>Si data Moderate decrease of the N=28 gap

- Shape coexistence in <sup>44</sup>S
- <sup>42</sup>Si very sensitive to details of the interaction
- Mean field (Gogny D1S, 5DCH), 2000 Strong decrease of the N=28 gap • <sup>44</sup>S: shape coexistence • <sup>42</sup>Si: very deformed (oblate: β=0.4)



S. Peru, M. Girod and J.F. Berger, EPJA9(2000)



### **Doubly magic <sup>42</sup>Si?**

#### nature

Vol 435|16 June 2005|doi:10.1038/nature03619

# LETTERS

# 'Magic' nucleus <sup>42</sup>Si

- □ <sup>42</sup>Si,<sup>43</sup>P produced from 1,2-proton removal
- Low statistics
- **□** In-beam γ spectroscopy
- □ SeGA Ge array (efficiency 2.5%)
- no γ observed in <sup>42</sup>Si interpreted as no bound excited state
- small 2-proton cross section interpreted as a sign of shell closure
- Conclusion of a closed shell <sup>42</sup>Si in contradiction with previous β decay lifetime S. Grévy *et al.* Phys. Lett. B **594** (2004)

**!!!** Wrong interpretation of the measurement **!!!** (see next)



## Cea Collapse of the N=28 shell closure

- <sup>42</sup>Si not magic... but very deformed
- GANIL experiment, LISE
- □ <sup>42</sup>Si produced from 1,2 proton removal at 39 A MeV
- $\Box$  in-beam  $\gamma$  with high photopeak efficiency (38%, Nal array « Château de crystal »)
- □ first 2<sup>+</sup> of <sup>42</sup>Si at low energy: 770(19) keV
- Results later confirmed with additional transitions (41<sup>+</sup> state ar 2173(14) keV) at RIBF
   S. Takeuchi *et al.*, Phys. Rev. Lett. **109**, 182501 (2012)



# Cea Shape coexistence in <sup>44</sup>S

□ transition between spherical and deformed ground-state configurations





### Mean field approaches

- prolate oblate shape coexistence
- □ Shell model approaches
  - prolate-spherical coexistence
  - prediction of a  $0^+_2$  state above the  $2^+_{0.5} \begin{vmatrix} 1 \\ 0^+_1 \\ 0^+_2 \end{vmatrix}$









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### Cell The merging of the N=20 & N=28 regions of deformation

- RIBF, RIKEN
- $\Box$  in-beam  $\gamma$  spectroscopy of <sup>34,36,38</sup>Mg
- Flat systematics of 2<sup>+</sup><sub>1</sub>,4<sup>+</sup><sub>1</sub>
   Low 2<sup>+</sup><sub>1</sub> energies (about 500 keV)
   R<sub>42</sub>>3, consistent with deforemd nuclei
- Merging of the N=20 island of inversion with the N=28 region of deformation

P. Doornenbal, Phys. Rev. Lett. 111, 212502 (2013)

Spectroscopy of <sup>40</sup>Mg measured in 2016 under analysis



### Second" island of inversion at N=40, an old story

The Physics around the doubly-magic <sup>78</sup>Ni Nucleus

From A. Poves

Leuven, Belgium November 4-5, 1996



in the intruder

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**Second**" island of inversion at N=40, an old story

Eur. Phys. J. A **16**, 55–61 (2003) DOI 10.1140/epja/i2002-10069-9

The European Physical Journal A

### New region of deformation in the neutron-rich <sup>60</sup><sub>24</sub>Cr<sub>36</sub> and <sup>62</sup><sub>24</sub>Cr<sub>38</sub>

O. Sorlin<sup>1,a</sup>, C. Donzaud<sup>1</sup>, F. Nowacki<sup>2</sup>, J.C. Angélique<sup>3</sup>, F. Azaiez<sup>1</sup>, C. Bourgeois<sup>1</sup>, V. Chiste<sup>1</sup>, Z. Dlouhy<sup>4</sup>, S. Grévy<sup>3</sup>, D. Guillemaud-Mueller<sup>1</sup>, F. Ibrahim<sup>1</sup>, K.-L Kratz<sup>5</sup>, M. Lewitowicz<sup>6</sup>, S.M. Lukyanov<sup>7</sup>, J. Mrasek<sup>4</sup>, Yu.-E. Penionzhkevich<sup>7</sup>, F. de Oliveira Santos<sup>6</sup>, B. Pfeiffer<sup>5</sup>, F. Pougheon<sup>1</sup>, A. Poves<sup>8</sup>, M.G. Saint-Laurent<sup>6</sup>, and M. Stanoiu<sup>6</sup>

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- <sup>7</sup> FLNR, JINR, 141980 Dubna, Moscow region, Russia
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**Abstract.** The neutron-rich nuclei  ${}^{60-63}_{23}$ V have been produced at GANIL via interactions of a 61.8 A · MeV  ${}^{76}$ Ge beam with a  ${}^{58}$ Ni target. Beta-decay to  ${}^{60-63}_{24}$ Cr has been investigated using combined  $\beta$ - and  $\gamma$ -ray spectroscopy. Half-lives of the  ${}^{60-63}$ V nuclei have been determined, and the existence of a beta-decay isomer in the  ${}^{60}$ V nucleus is strongly supported. The observation of low-energy  $2^+$  states in  ${}^{60}$ Cr (646 keV) and  ${}^{62}$ Cr (446 keV) suggests that these isotopes are strongly deformed with  $\beta_2 \sim 0.3$ . This is confirmed by shell model calculations which show the dominant influence of the intruder g and d orbitals to obtain low  $2^+$  energies in the neutron-rich Cr isotopes.

### Ce2 Fe and Cr isotopes



#### PHYSICAL REVIEW C 81, 051304(R) (2010)

RAPID COMMUNICATION

### Collectivity at N = 40 in neutron-rich <sup>64</sup>Cr

A. Gade,<sup>1,2</sup> R. V. F. Janssens,<sup>3</sup> T. Baugher,<sup>1,2</sup> D. Bazin,<sup>1</sup> B. A. Brown,<sup>1,2</sup> M. P. Carpenter,<sup>3</sup> C. J. Chiara,<sup>3,4</sup> A. N. Deacon,<sup>5</sup>
S. J. Freeman,<sup>5</sup> G. F. Grinyer,<sup>1</sup> C. R. Hoffman,<sup>3</sup> B. P. Kay,<sup>3</sup> F. G. Kondev,<sup>6</sup> T. Lauritsen,<sup>3</sup> S. McDaniel,<sup>1,2</sup> K. Meierbachtol,<sup>1,7</sup> A. Ratkiewicz,<sup>1,2</sup> S. R. Stroberg,<sup>1,2</sup> K. A. Walsh,<sup>1,2</sup> D. Weisshaar,<sup>1</sup> R. Winkler,<sup>1</sup> and S. Zhu<sup>3</sup>

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<sup>2</sup>Department of Physics and Astronomy, Michigan State University, East Lansing, Michigan 48824, USA





# Cell Fe and Cr isotopes

GANIL

#### PHYSICAL REVIEW C 81, 061301(R) (2010)

#### Onset of collectivity in neutron-rich Fe isotopes: Toward a new island of inversion?

J. Ljungvall,<sup>1,2,3</sup> A. Görgen,<sup>1</sup> A. Obertelli,<sup>1</sup> W. Korten,<sup>1</sup> E. Clément,<sup>2</sup> G. de France,<sup>2</sup> A. Bürger,<sup>4</sup> J.-P. Delaroche,<sup>5</sup> A. Dewald,<sup>6</sup>



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**RAPID COMMUNICATION** 

# Cea A new Island of inversion?



- □ Reduction of the fp-gds gap towards the neutron dripline
- □ strong similarities to the N=20 shell gap behavior

Difference between shell-model and mean-field for the neutron spherical shell gap at N=40

L. Gaudefroy et al., Phys. Rev. C 80, 064313 (2009)



# Cera Shell model description of N=40 isotones



S.M. Lenzi et al., Phys. Rev. C 82, 054301 (2010).

Slide by F. Nowacki, IPHC

# **Cerror The SEASTAR program at the RIBF**

- RIBF is the laboratory where the spectroscopy of the most exotic species can be accessed ٠
  - Dedicated program of search for new 2<sup>+</sup> states: experimental campaigns in 2014, 215 and 2017



DALI2 + thick liquid hydrogen target (MINOS)

•

•

# Cera Features of in-beam gamma spectroscopy









Program based on (p,2p), (p,pn), (p,3p)

A. Obertelli et al., Eur. Phys. Jour. A 50, 8 (2014)

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# Ce2 Fe and Cr isotopes



### « Second » island of inversion



C. Santamaria et al., Phys. Rev. Lett. 115, 192501 (2015).

# Cera Extension of the N=40 Iol towards N=50



□ Full *fp* for protons, full *sdg* for neutrons

- □ First 2<sup>+</sup> excited state predicted to be intruder state
- Prediction of dissapearance of N=50 shell closure below <sup>78</sup>Ni

F. Nowacki et al., Phys. Rev. Lett. 117, 272501 (2016)

# Cera N=32 and N=34: shell closures ?

- □ Neutron-rich *fp* shell
- □ Attractive interaction between  $\pi 1f_{7/2}$  and  $\nu 1f_{5/2}$  orbitals is important; responsible for some features of nuclear shell evolution in this mass region

T. Otsuka et al., Phys. Rev. Lett. 95 (2005) 232502

□ As protons are removed from the  $\pi f_{7/2}$  orbital (i.e., from <sub>28</sub>Ni to <sub>20</sub>Ca) the strength of the  $\pi$ v interaction weakens, causing the  $v f_{5/2}$  orbital to shift up in energy relative to  $v p_{1/2}$  and  $v p_{3/2}$ 



# Cea Spectroscopy of <sup>54</sup>Ca

<sup>56</sup>Ti, <sup>55</sup>Sc + Be -> <sup>54</sup>Ca + X at the RIBF (DALI2)
<sup>70</sup>Zn primary beam (100 pnA max)
<sup>56</sup>Ti 120 pps/pnA, <sup>55</sup>Sc 12 pps/pnA



#### D. Steppenbeck et al., Nature 502 (2013)



# Cerror Shell model calculations for 54Ca



# Cerror Beyond mean-field calculations for neutron-rich Ca

- overall proper predictions from GCM with Gogny D1S
- □ cranking necessary for a proper prediction of absolute excitation energies



T. R. Rodriguez, Phys. Rev. Lett. 99, 06201 (2007)





- The shell model and shell closures
- Origins of shell evolution
  - diffusiveness of the nuclear surface
  - spin-isospin and tensor terms of the NN interaction
  - reduction of spin-orbit
  - 3N forces
- RI beam production and future facilities
- The N=16 "new magic number", collapse of N=20, 28 shell closures
- Mirror region: the N=40 island of inversion and the <sup>78</sup>Ni region
- Are N=32 and 34 "new" magic numbers?
- Heavier doubly-magic nuclei: <sup>100</sup>Sn and <sup>132</sup>Sn

# Cellectivity in the <sup>100</sup>Sn region



# Ce2 Collectivity in the <sup>100</sup>Sn region

- □ Gamow-Teller transition change proton into neutron
- Experimental probes: charge exchange reactions, beta-decay, EC decay
- **Gamow-Teller strength**  $B_{GT} \sim 1/ft$
- Gamow-Teller strenght sensitive to shell closure



Hinke, PhD Thesis, Univ. Munchen (2010) Hinke, Nature 486, 341 (2012)



□ All GT strenght on one final state  $^{100}$ In(1<sup>+</sup>)

 $\square$  B<sub>GT</sub>=9.1 exceptionally large

### I strong shell closure at <sup>100</sup>Sn



### Cell Structure of <sup>132</sup>Sn

- □ Holifield RIB facility at Oak Ridge: re-accelerated fission fragment beam, 4.5 MeV/nucleon
- Machine is decommissionned today
- $\Box$  <sup>132</sup>Sn(d,p)<sup>133</sup>Sn via missing mass
- Single-particle nature of neutron states in <sup>133</sup>Sn confirmed. Validates <sup>132</sup>Sn as magic [Neutron transfer: see tomorrow's lecture]



Q-value spectrum for the  $^{132}$ Sn(d,p) $^{133}$ Sn reaction at 54° in the centre of mass.





- □ The picture of the nuclear shell structure has been explored since 40 years
- □ The nuclear shell model offers a succesfull framework to interpretate observed structure changes
- □ EDF-based approaches give overall sucessfull predictions
- □ No breakdown of first-developped models
- In the shell model picture, monopole drifts (spin-orbit,spin-isospin, tensor) develop and modify « spherical mean-field » gaps, while the interplay with correlations (pairing, quadrupole) drives the physics
- **Three-body forces** necessary for a quantitative understanding of shell structure and evolution
- Disappearance of N=20, 28 as shell closures in the neutron-rich region
- Joint appearance of N=16 as a new magic number
- N=32,34 intepreted as new shell closure (at least show subshell effects)
- Island of inversion at N=20, N=40 (mirror phenomena)
- prediction of disappearance of the N=50 shell closure below <sup>78</sup>Ni
- <sup>100</sup>Sn (N=Z=50) (unknown spectroscopy yet) and <sup>132</sup>Sn (N=82) seem doubly closed shell nuclei
- □ Shape / configuration coexistence exist everywhere across the nuclear landscape.